

# Identifying the QPO frequencies of GRS 1915+105 as General Relativistic Dynamic ones

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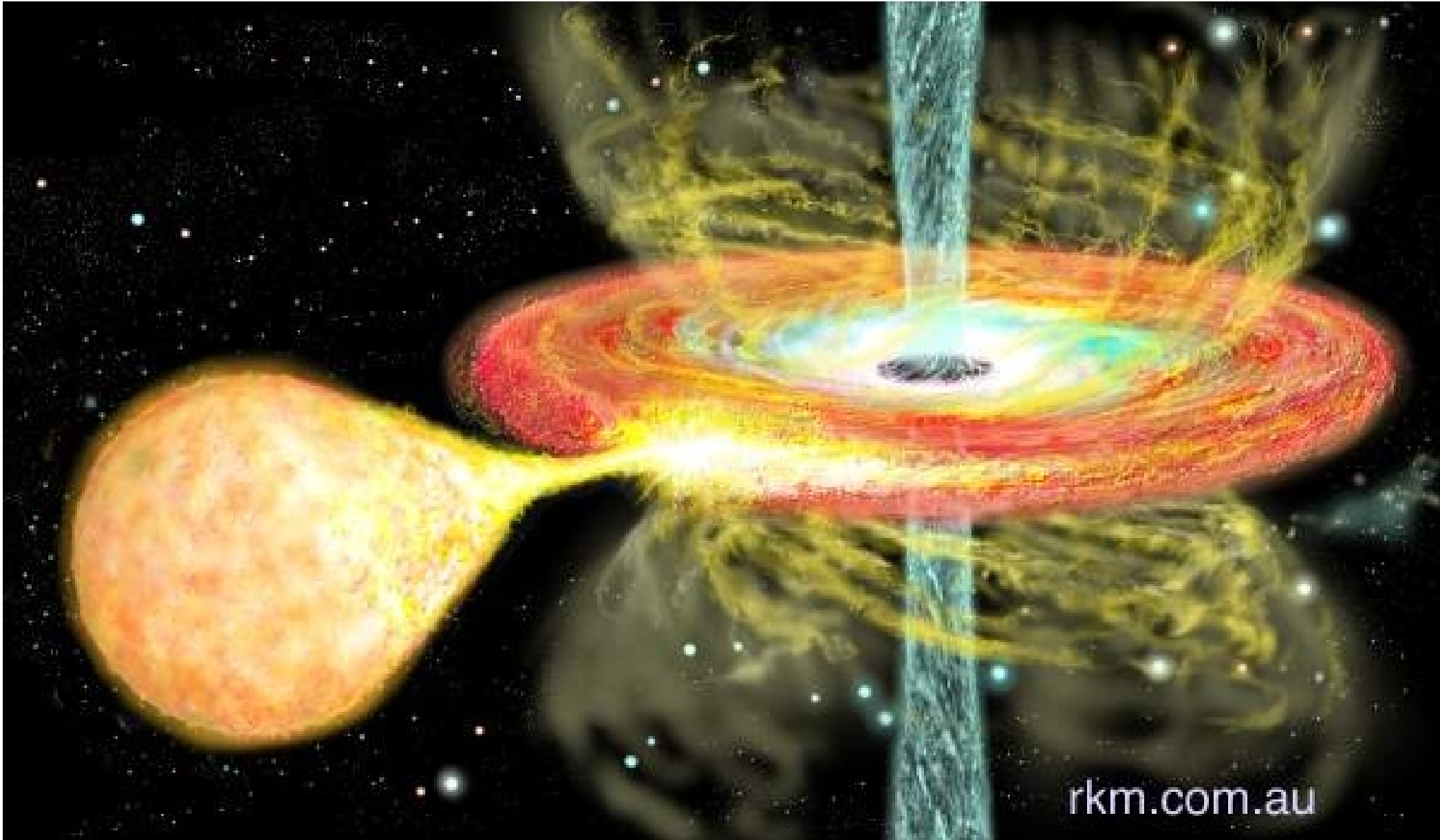
**Collaborators: Divya Rawat, Honghui Liu,  
Ruchika Dhaka**

**and J.S. Yadav, P. Jain, Ji, Long; C. Bambi, Y. Zhang**

**Based On: Misra, Rawat et al. 2020, Liu, H. et al. 2020, Dhaka  
et al. (in preparation) and Rawat et. al. (in preparation)**



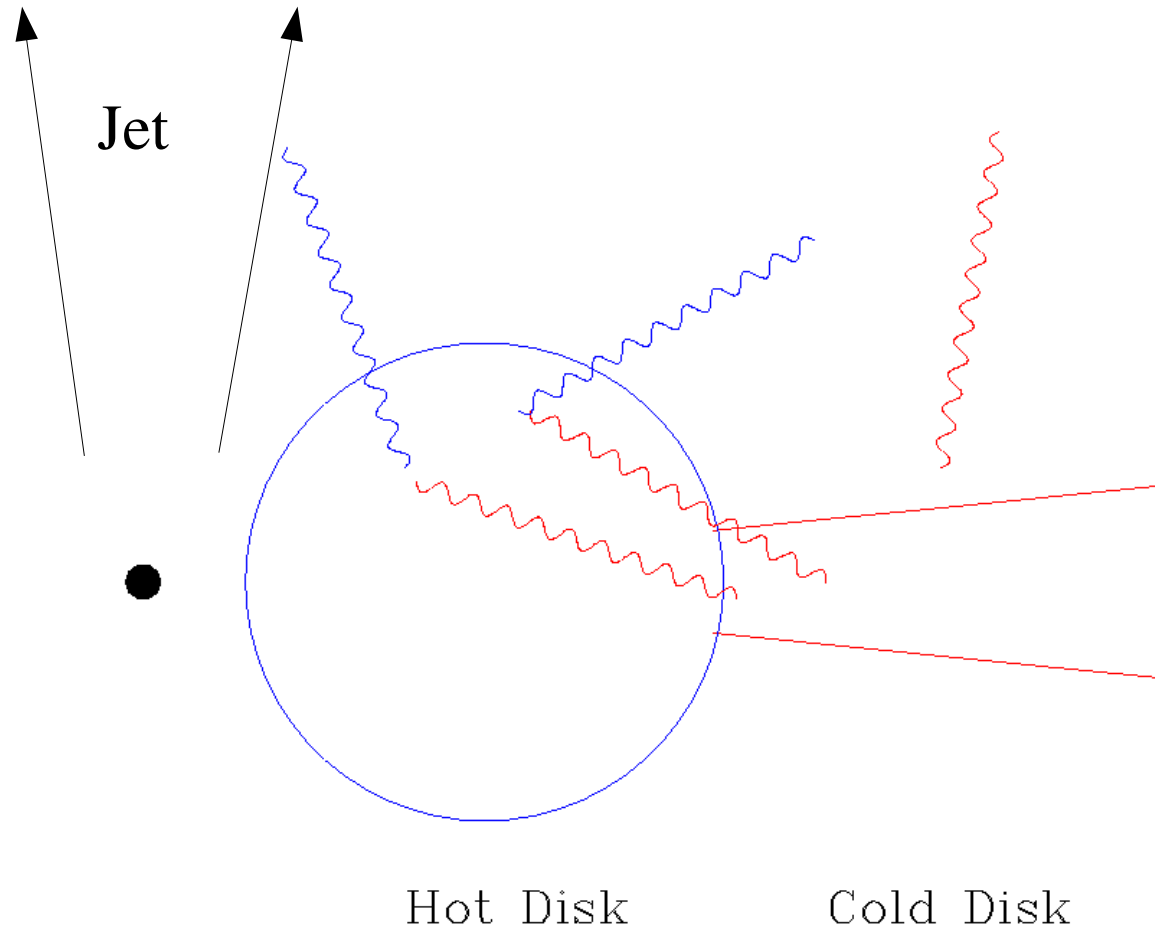
# *Picture of a Black Hole Binary*



rkm.com.au

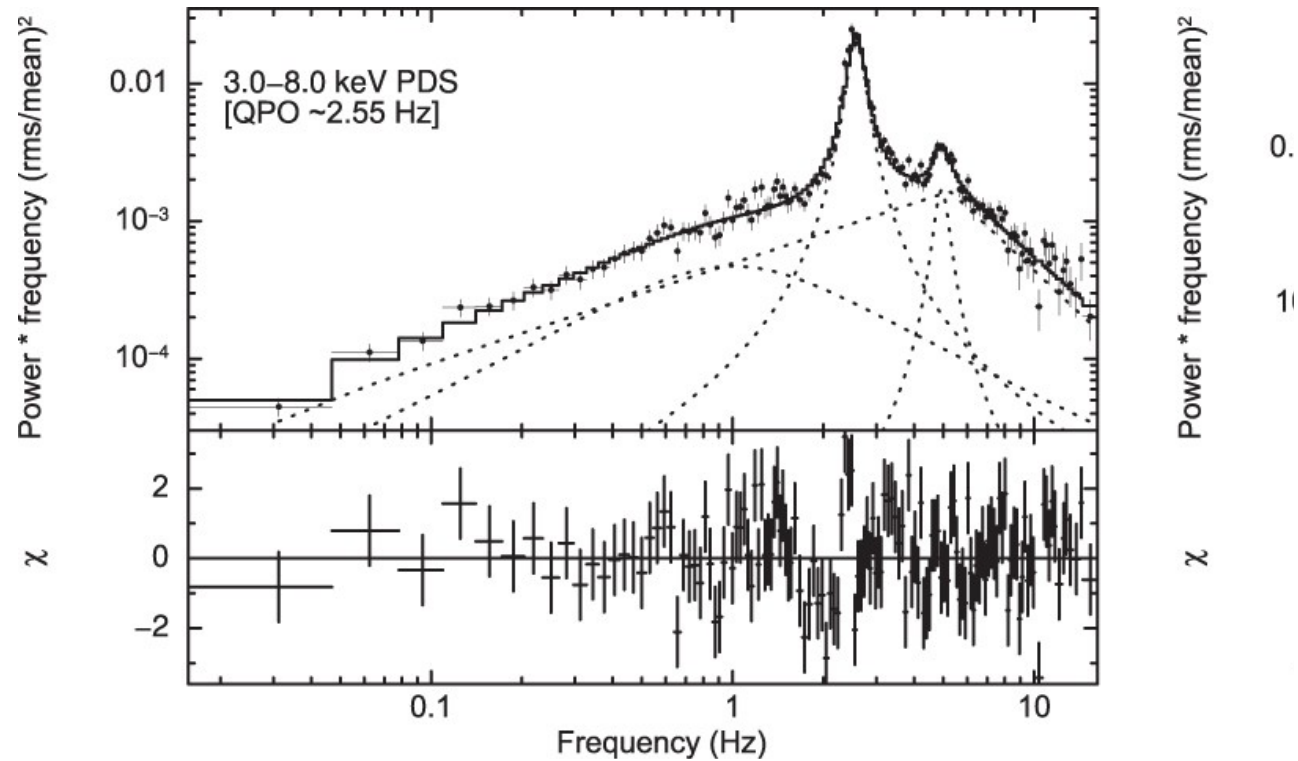


# *What we know about Black Hole Binaries...*





# *AstroSAT/LAXPC: GRS 1915+105, QPO*



Quasi Periodic Oscillations (QPO)

Yadav et al. 2016



# *Origin of QPOs...*

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- Identification with characteristic frequency of the system...
- Relativistic test particle frequencies
  - ➔ Keplerian Frequency
  - ➔ Epicyclic Frequency
  - ➔ Lenz-Thirring Frequencies

These frequencies depend only on radii



# *Origin of QPOs...*

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- Characteristic time-scales of the accretion disk
  - ➔ The sound crossing time,  $R/c_s$
  - ➔ The viscous time-scale,  $R/\nu_r$

These frequencies depend on radius and accretion rate



- *“Black holes in binary systems: Observational appearance”*. **Shakura, N. I. & Sunyaev, R. A.** 1973 *A&A*, 24, 337 - 355 **(18 Pages)**
- *“Astrophysics of black holes”*, **Novikov, I. D & Thorne, K. S.** 1974, *Black holes (Les astres occlus)*, p. 343-450 **(107 pages)**



# *Origin of QPOs...*

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Relativistic test particle frequencies

- ➔ Keplerian Frequency
- ➔ Epicyclic Frequency
- ➔ Lenz-Thirring Frequency

Characteristic time-scales of the accretion disk

- ➔ The sound crossing time,  $R/c_s$
- ➔ The viscous time-scale,  $R/\nu_r$





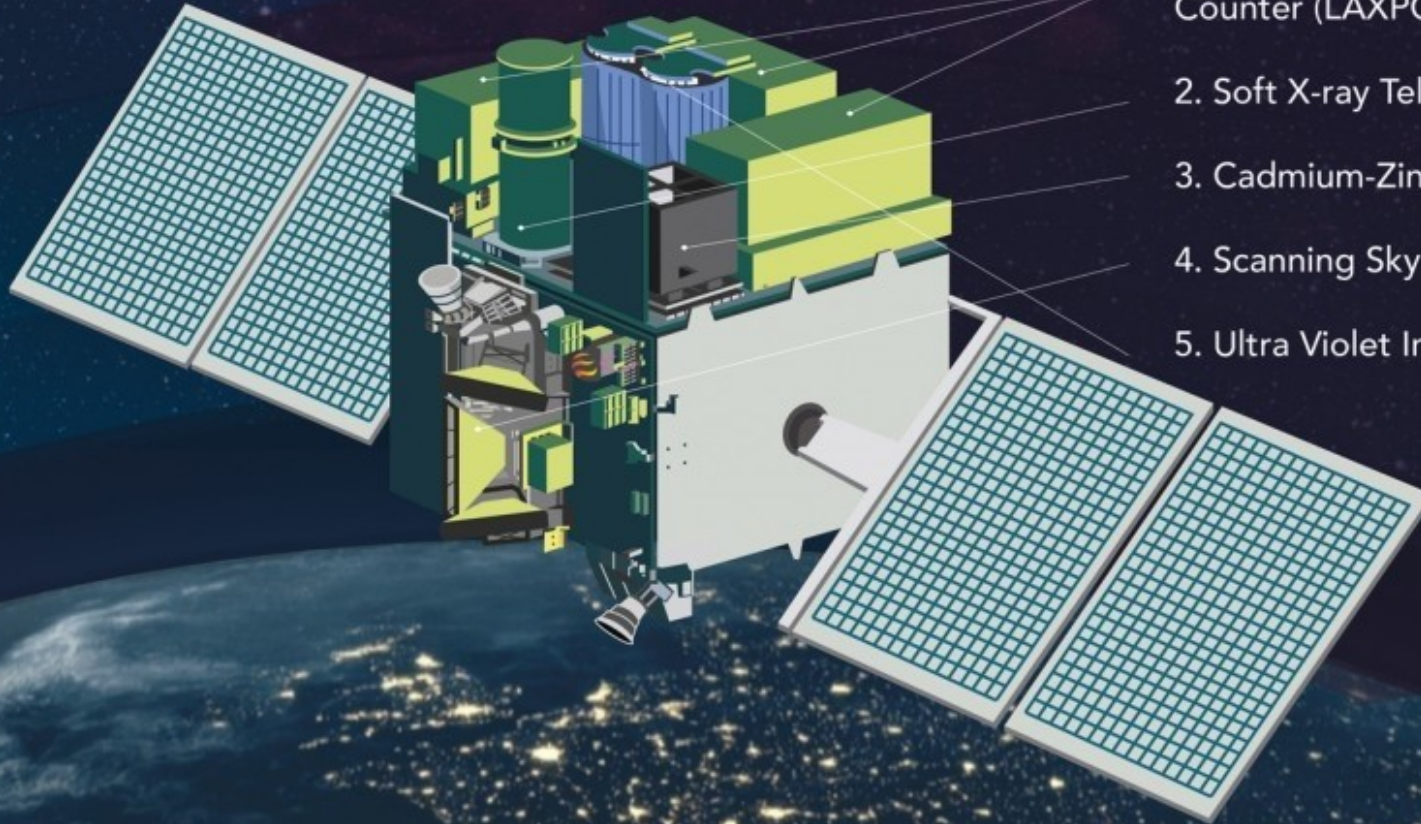
# *AstroSat*

India's first Multiwavelength Space Observatory

## ASTROSAT

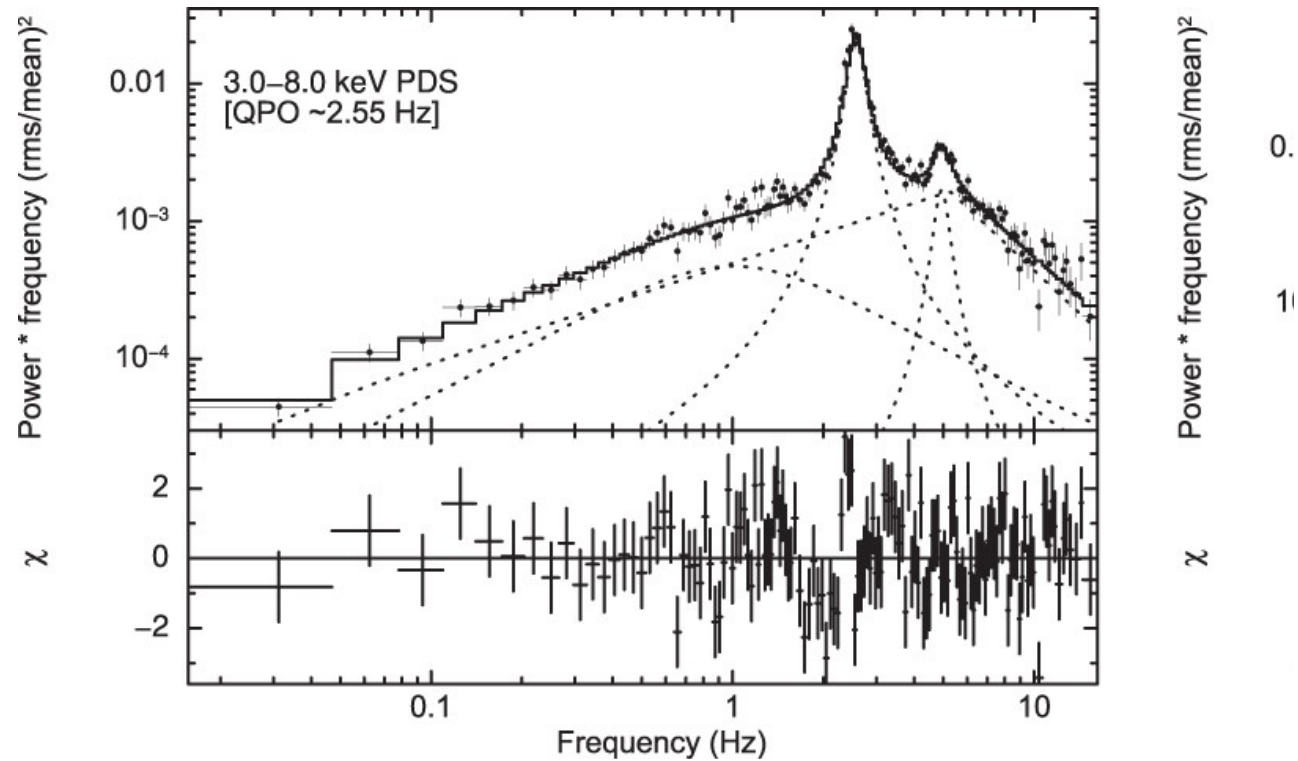
The 5 telescopes of the Astrosat

1. Large Area X-ray Proportional Counter (LAXPC)
2. Soft X-ray Telescope (SXT)
3. Cadmium-Zinc-Telluride Imager (CZTI)
4. Scanning Sky Monitor (SSM)
5. Ultra Violet Imaging Telescope (UVIT)





# AstroSAT/LAXPC: GRS 1915+105, QPO

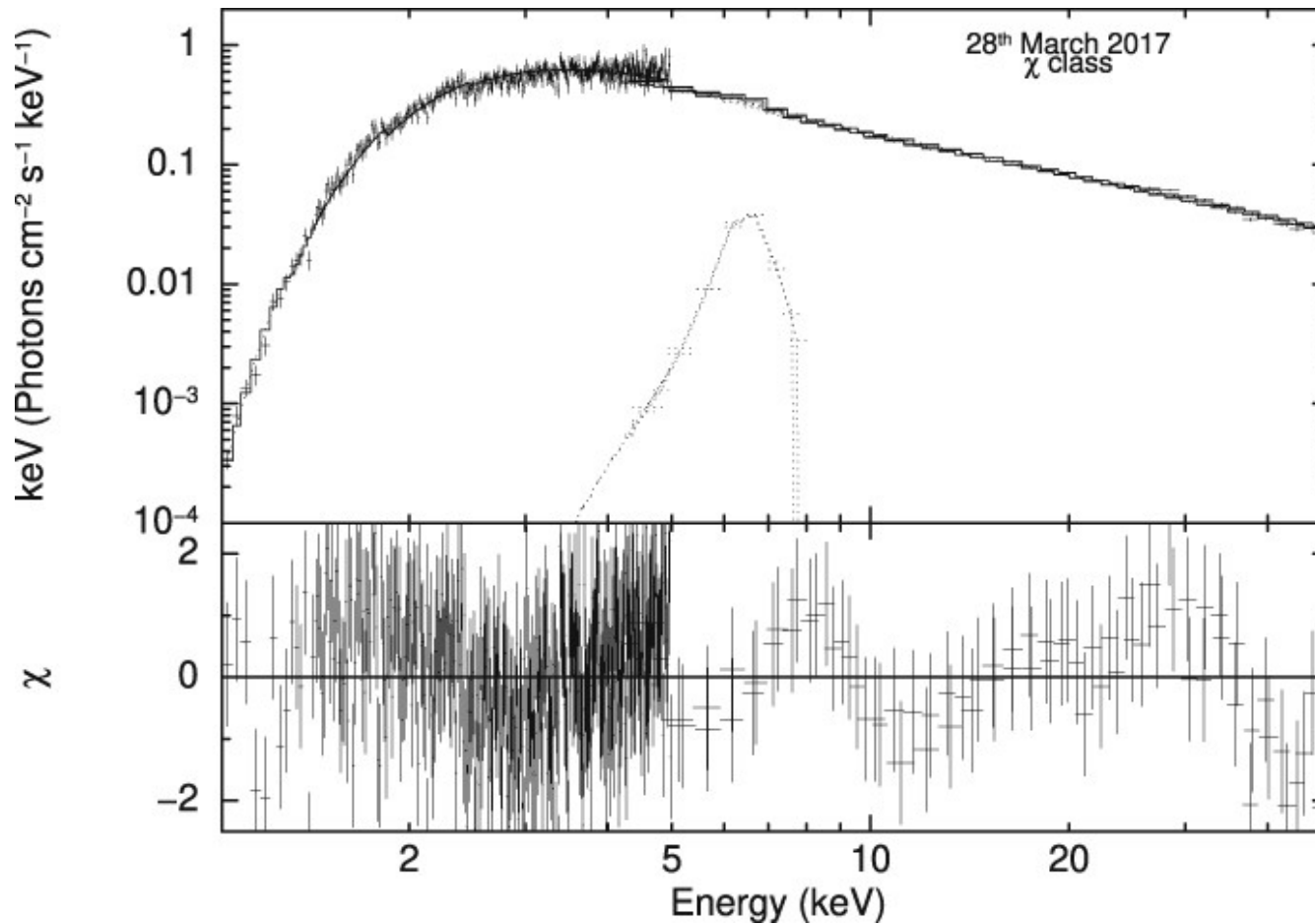


Quasi Periodic Oscillations (QPO)

Yadav et al. 2016



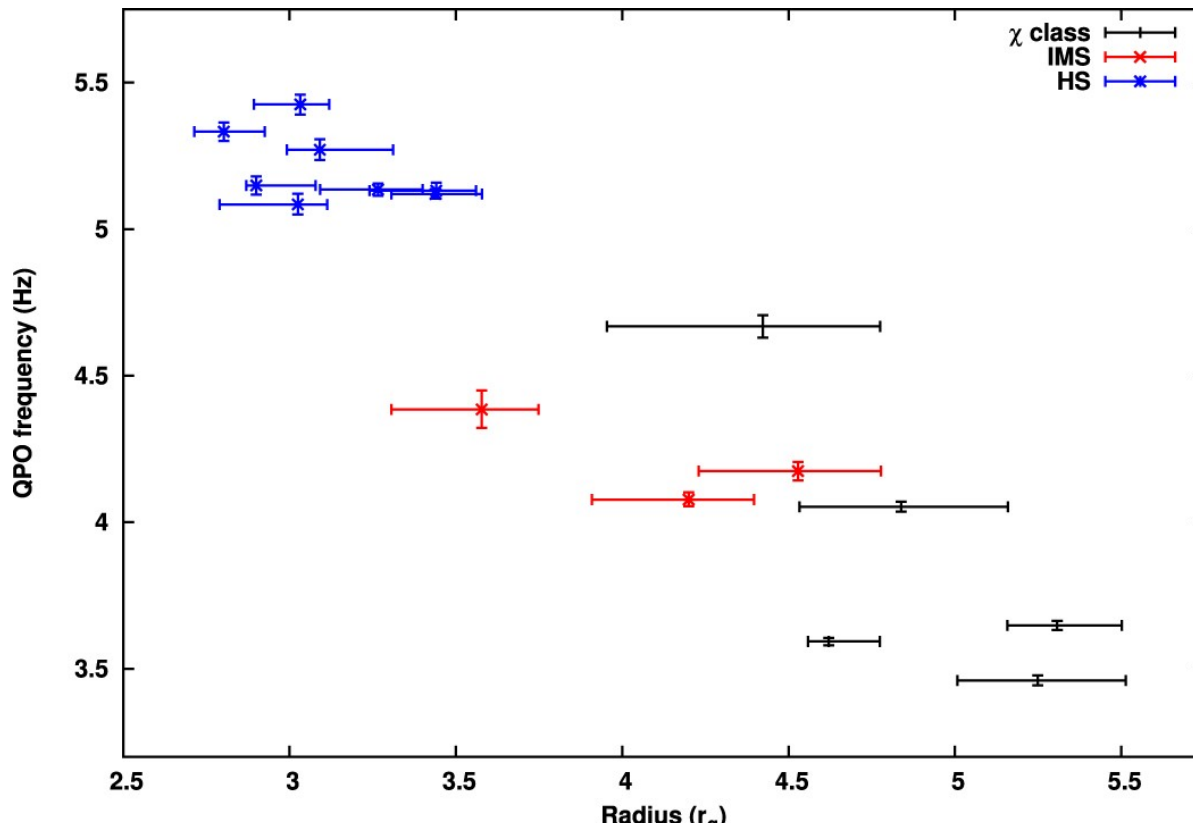
# *AstroSat/LAXPC & SXT Spectra*



Spectra fitting with AstroSat--→ **Estimate both inner disk radius and Accretion rate**



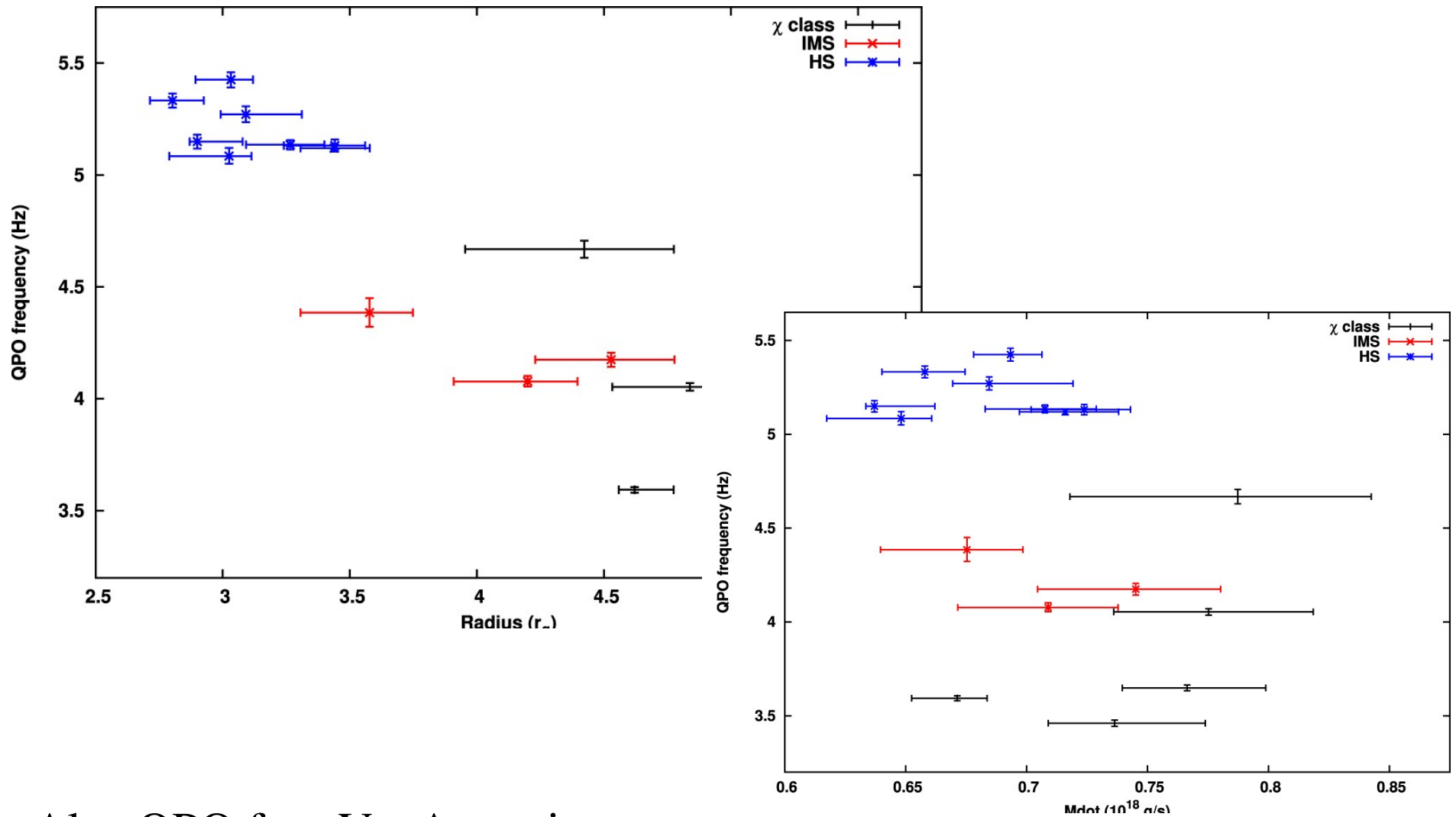
# QPO Frequency and inner disk radius



**Now we can see how the QPO Frequency varies with inner disk radius**



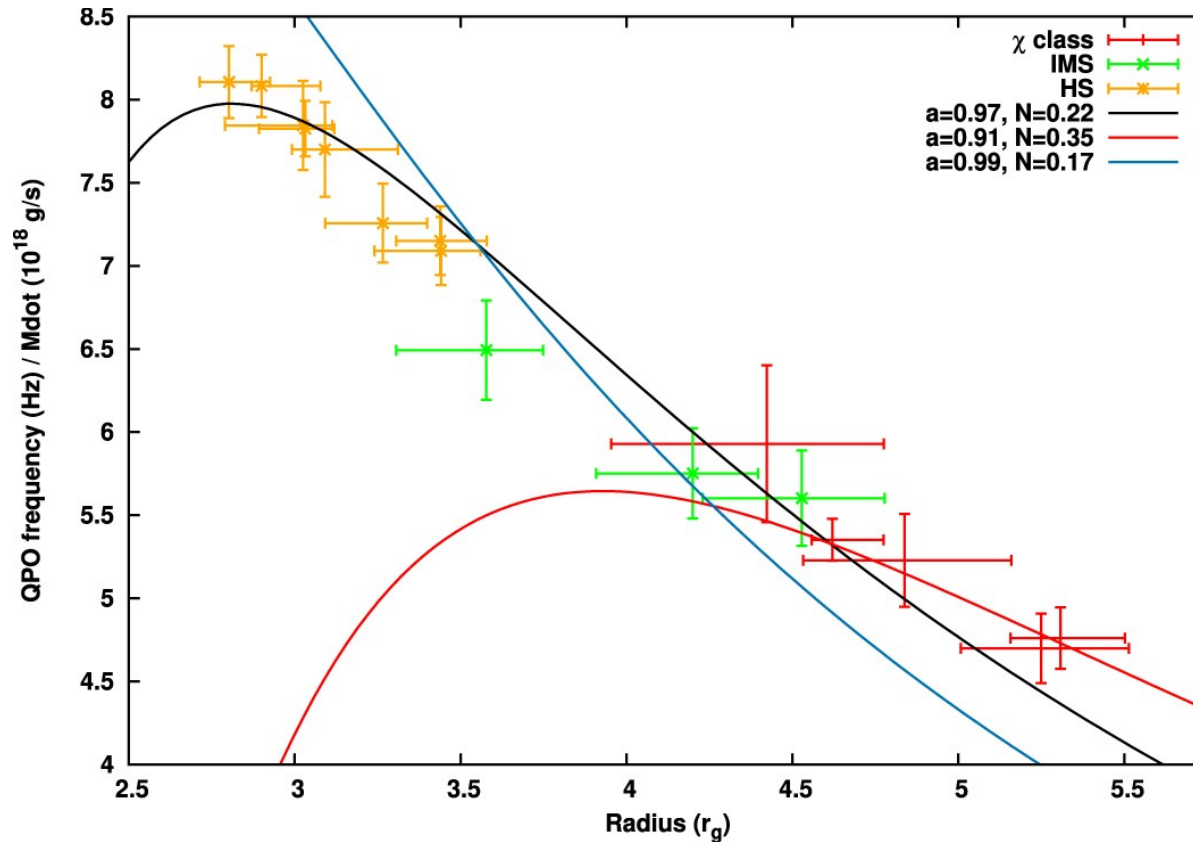
# QPO frequency and inner disk radius



Also QPO freq Vs. Accretion rate....



# QPO freq and disk radius and accretion rate



QPO Frequency DIVIDED by Accretion rate is much better correlated with inner disk radius...

Misra, Rawat, Yadav & Jain 2020



# *A General Relativity prediction....*

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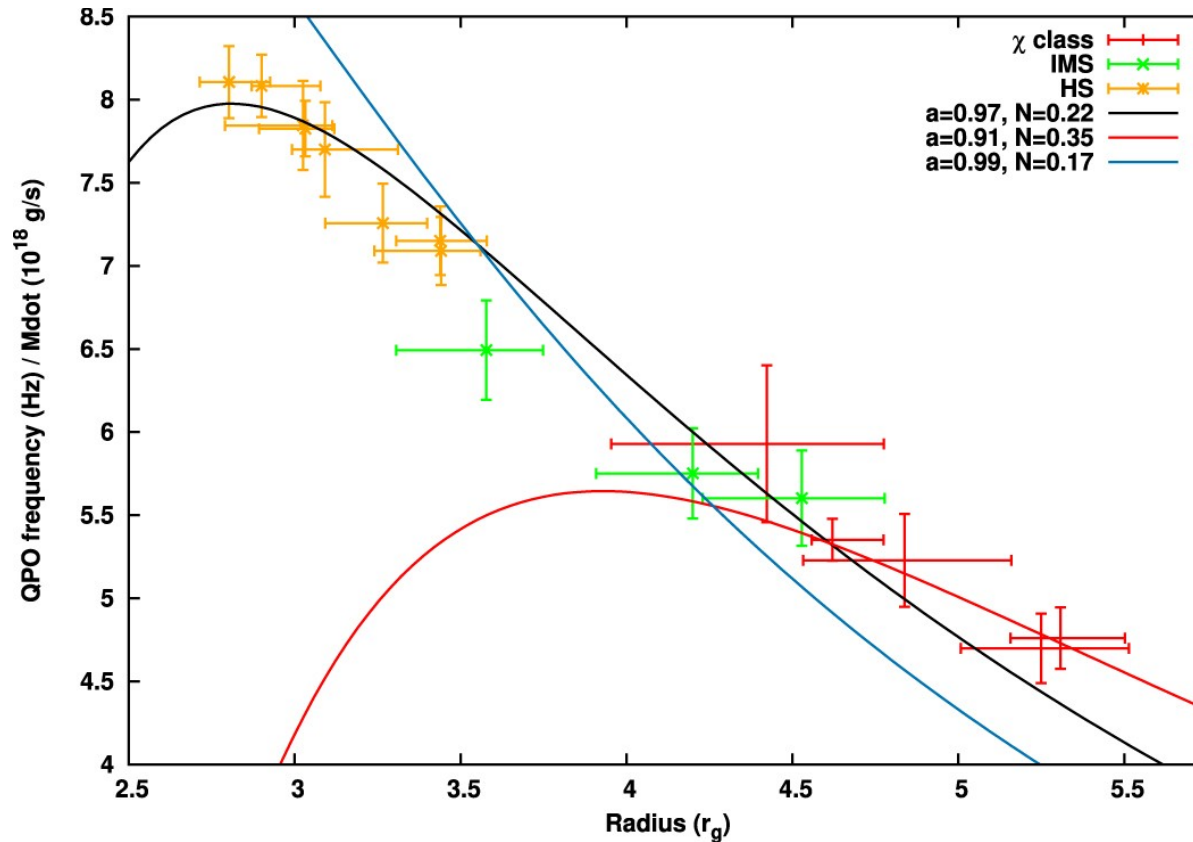
Novikov and Thorne derived that the Dynamic frequency (inverse of the sound crossing time) at a radius is given by...

$$\frac{f_{dyn}}{\dot{M}_{18}} = N \, 8979 \, \text{Hz} \, (r/r_g)^{-2.5} \, (M/12.4M_{\odot})^{-2} \\ \times A^1 B^{-2} D^{-0.5} E^{-0.5} L$$

Frequency divided by the accretion rate is a function of only radius



# A General Relativity prediction....



The black line is the predicted behavior for a black hole with Spin parameter = 0.97

Misra, Rawat, Yadav & Jain 2020

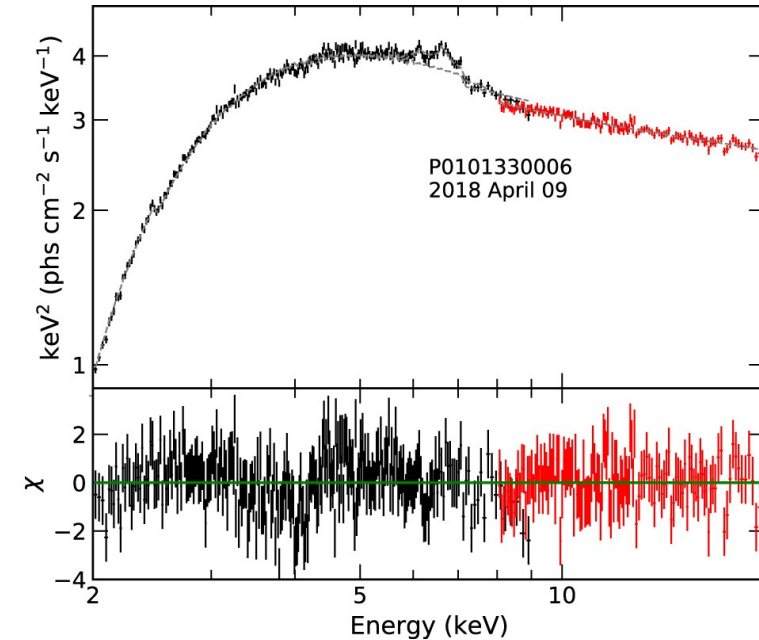
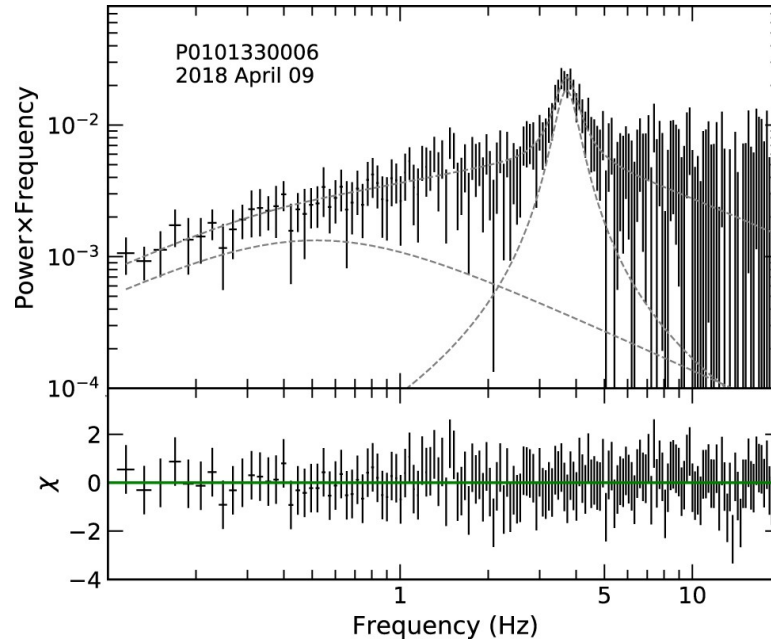




# *Insight HXMT*

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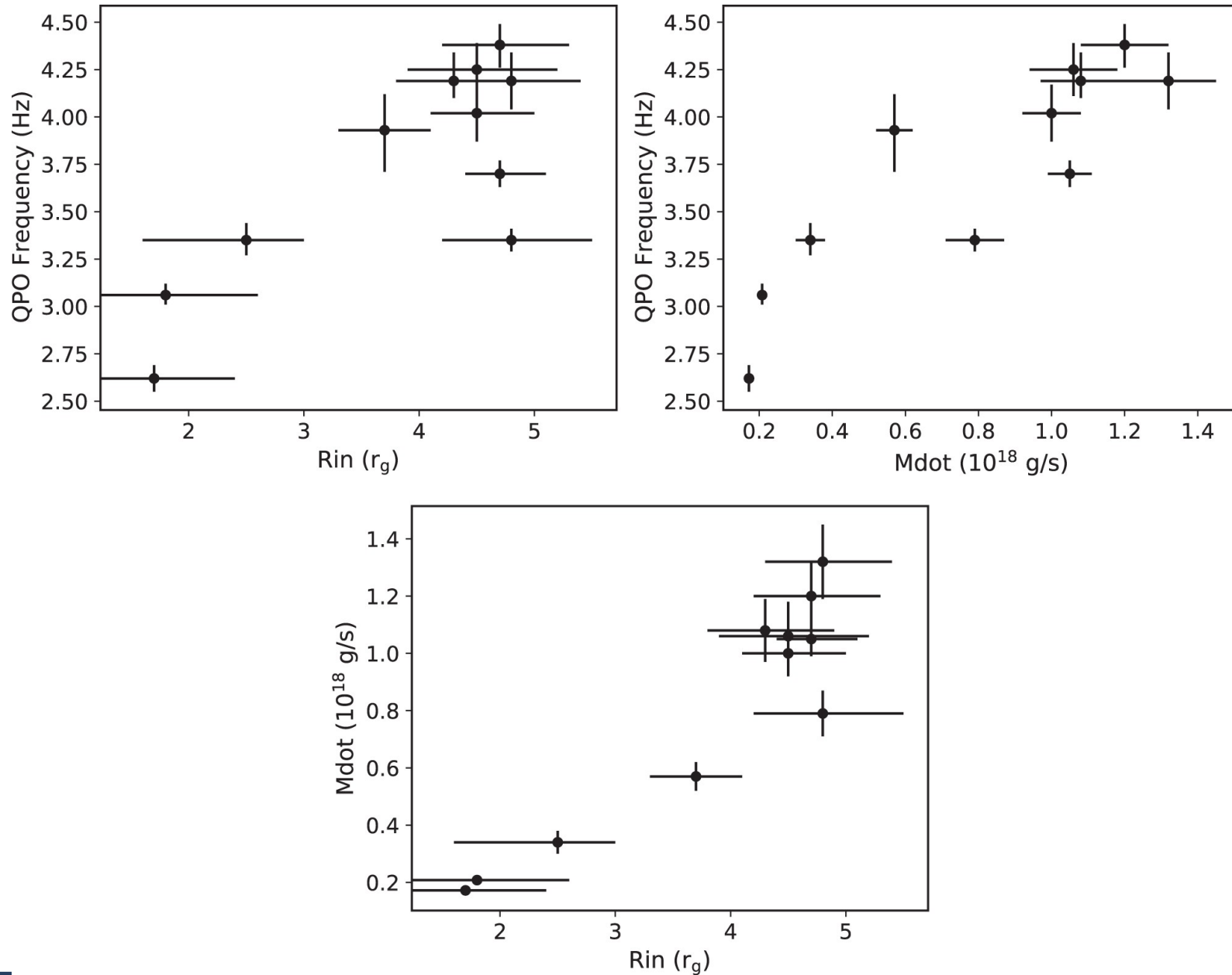




Spectra fitting with Insight-HXMT--  
→ Estimate both inner disk radius and Accretion rate



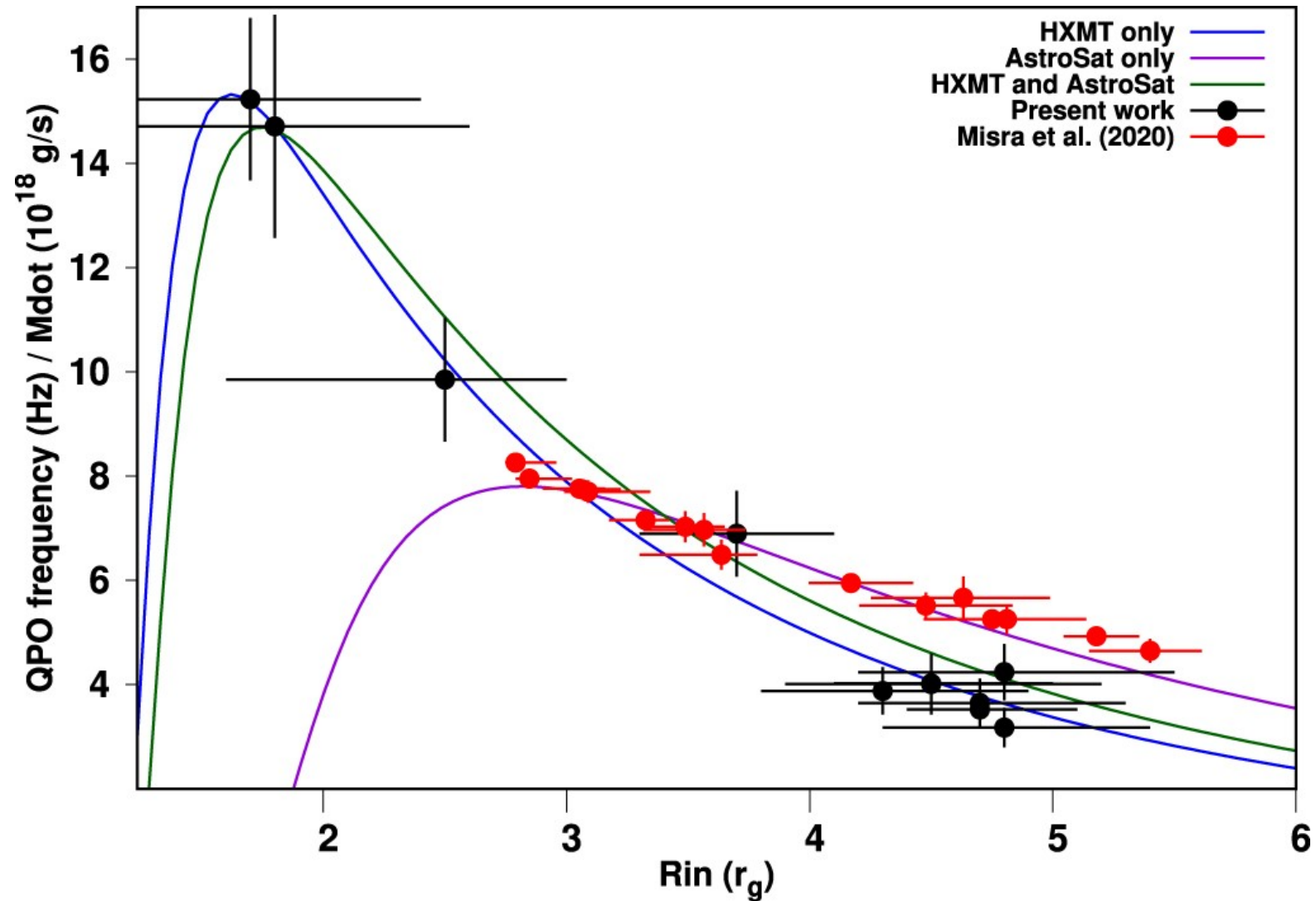
# Insight-HXMT



**Growing black holes: accretion and mergers**  
**Kathmandu, May 2022**

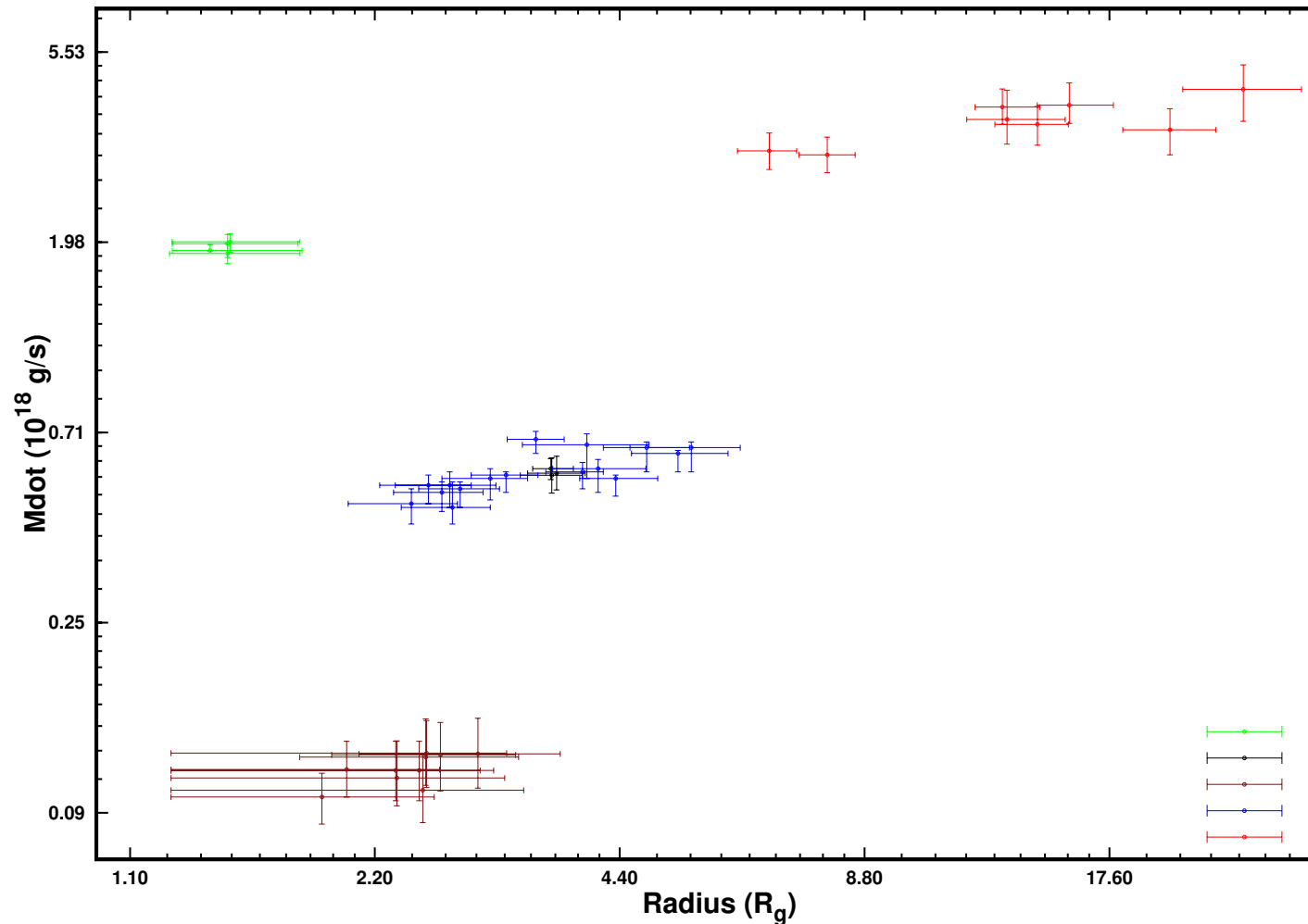


# Insight-HXMT





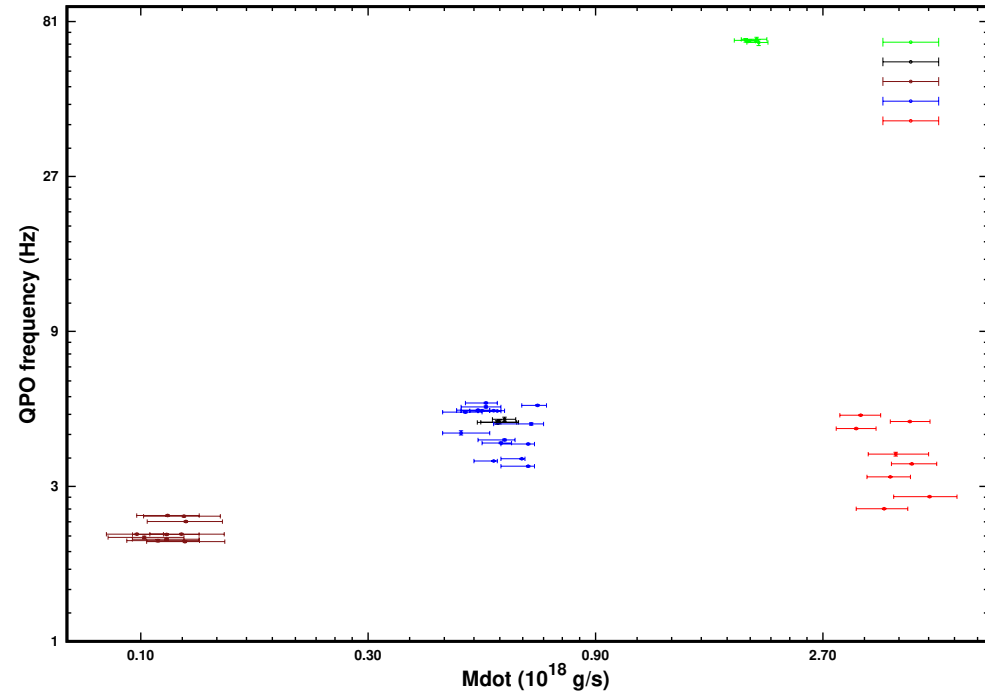
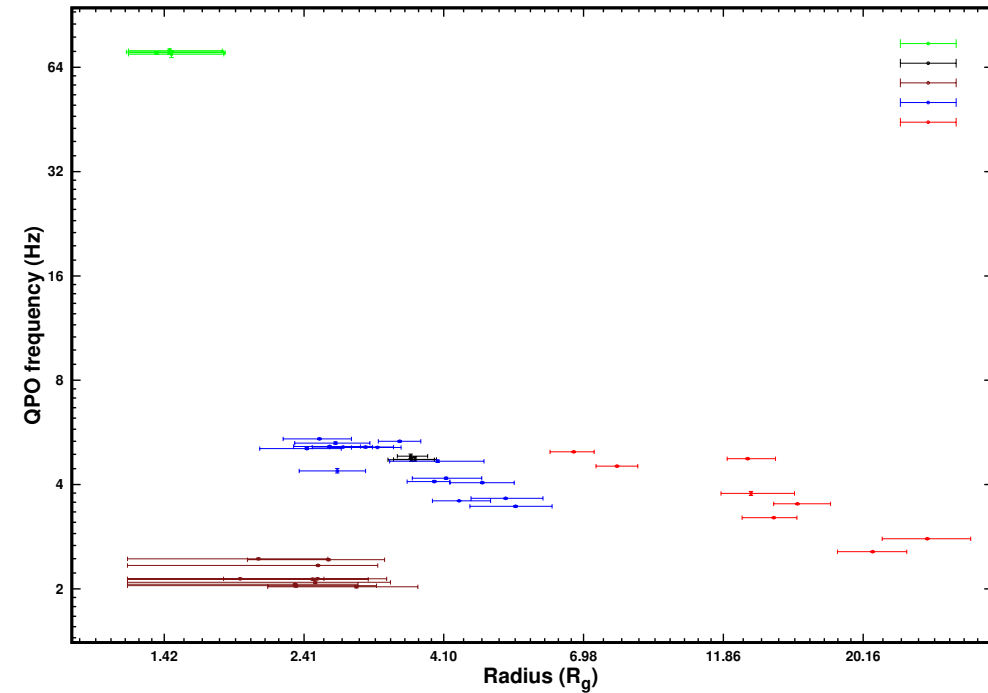
# More AstroSat Observations: PRELIMINARY RESULTS



Ruchika Dhaka et. al. in preparation



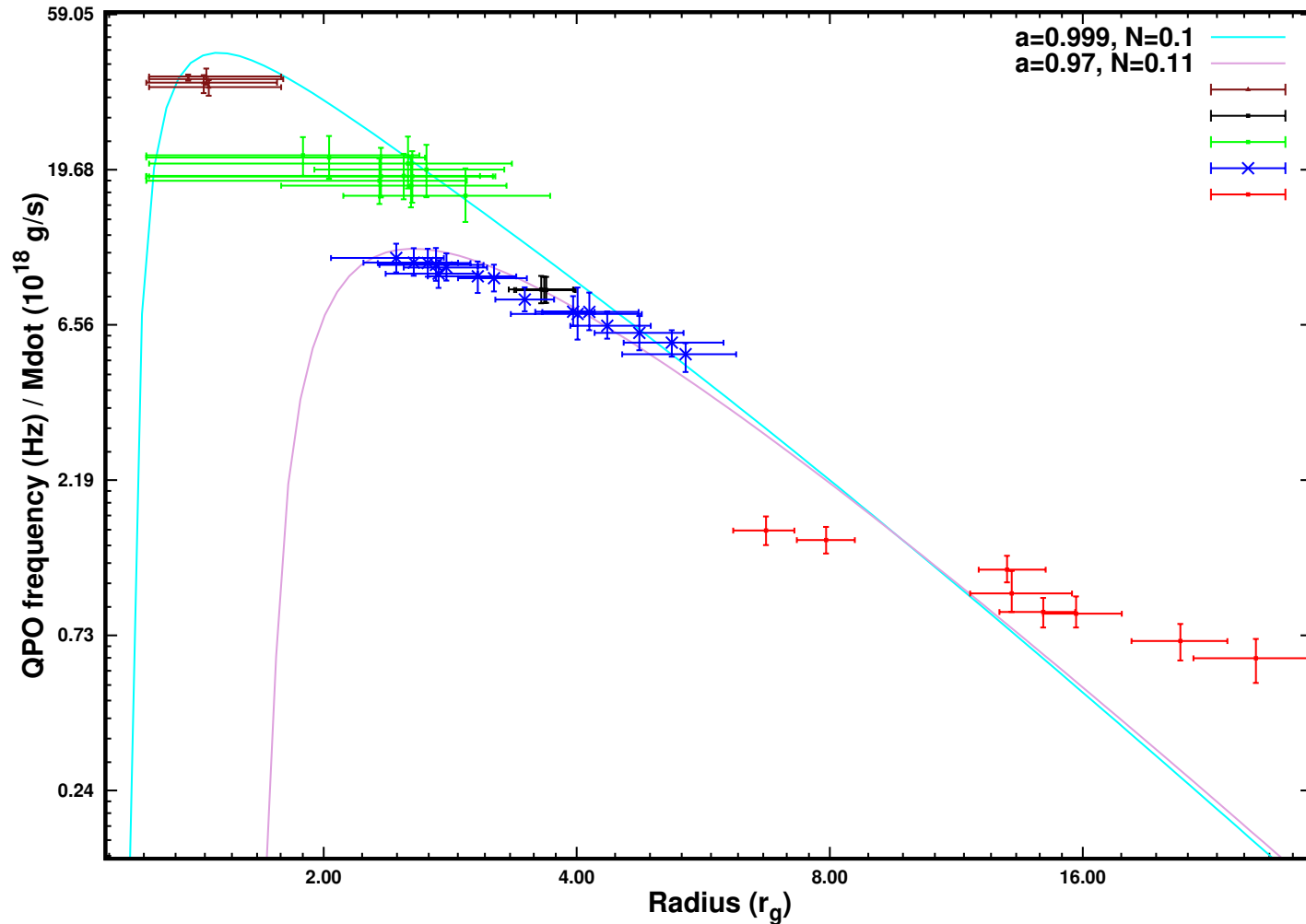
# More AstroSat Observations: PRELIMINARY RESULTS



Ruchika Dhaka et. al. in preparation



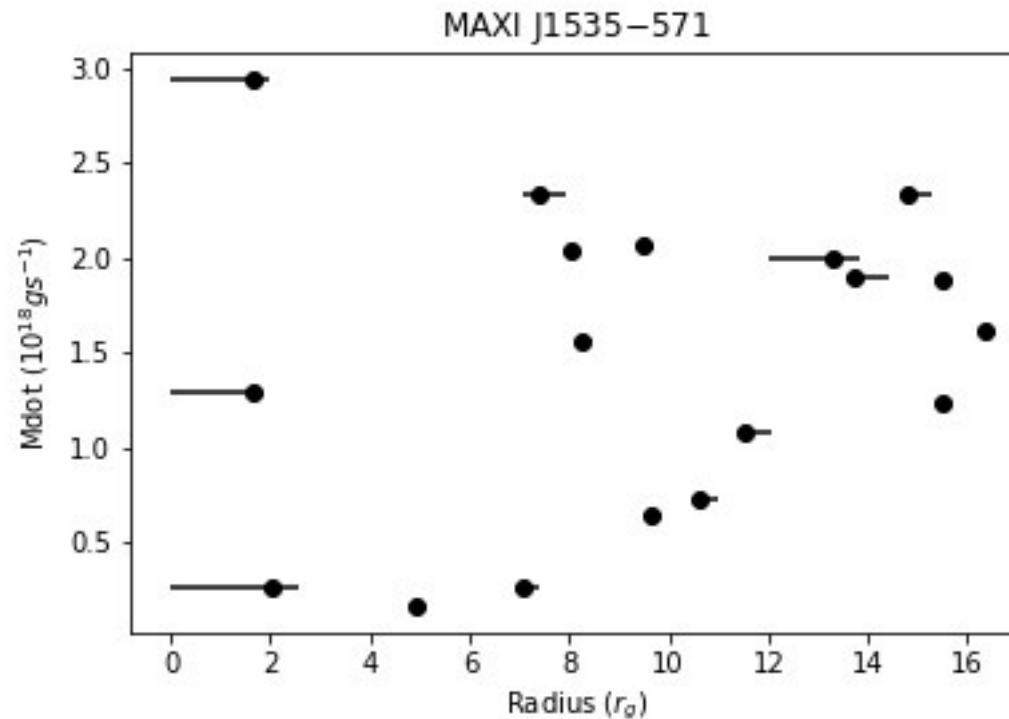
# More AstroSat Observations: PRELIMINARY RESULTS



Ruchika Dhaka et. al. in preparation



# NICER Observations: PRELIMINARY RESULTS

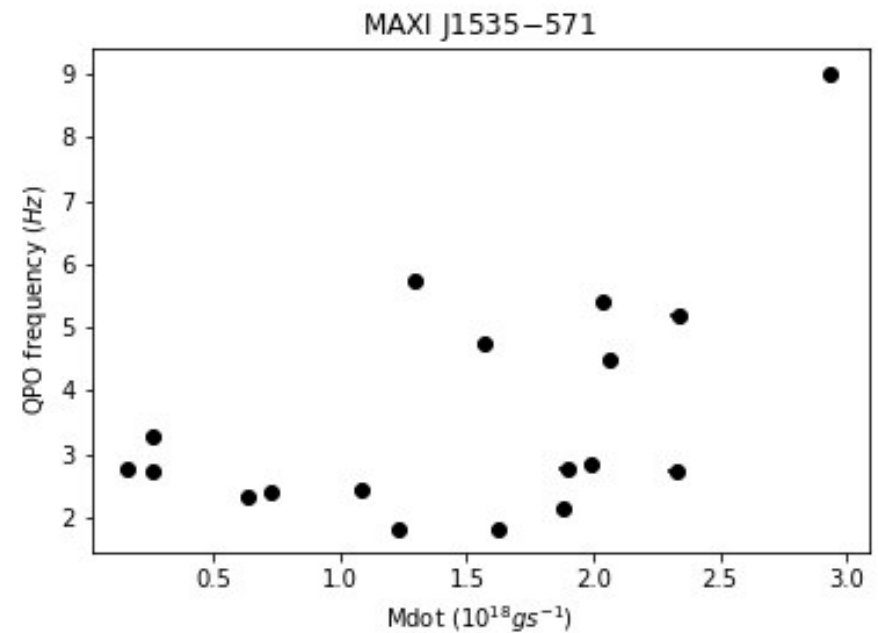
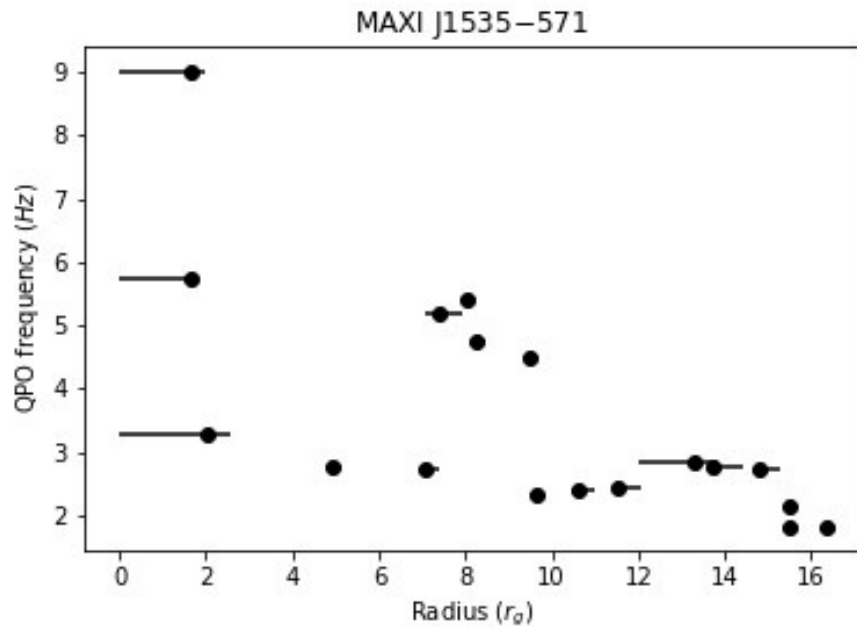


Divya Rawat et. al. in preparation





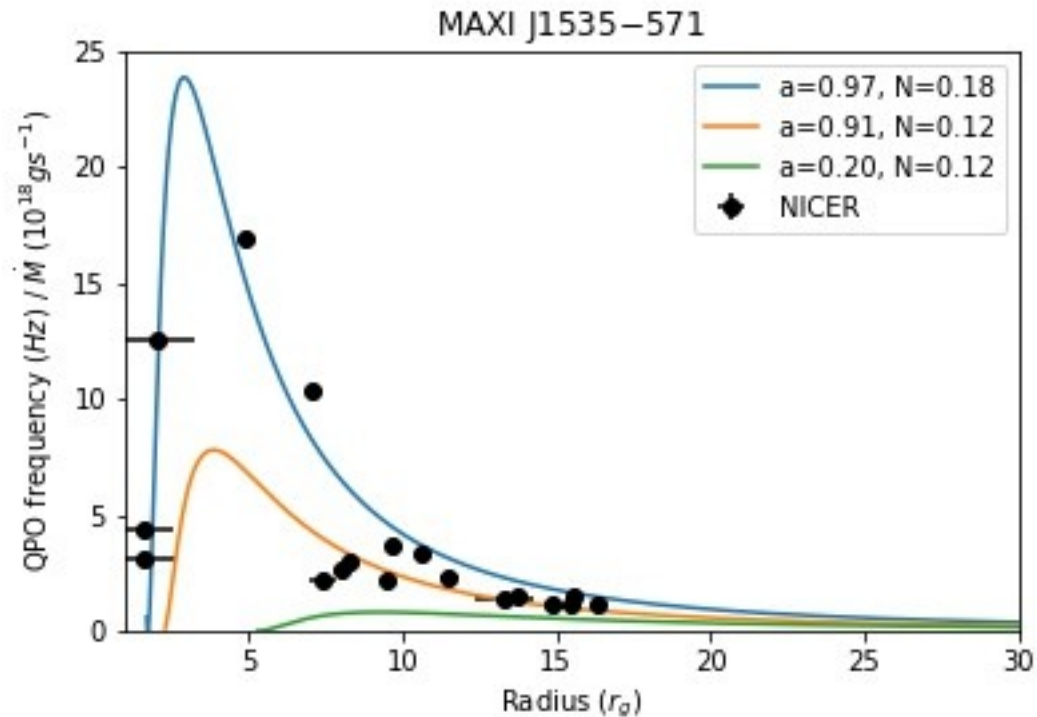
# NICER Observations: PRELIMINARY RESULTS



Divya Rawat et. al. in preparation



# NICER Observations: PRELIMINARY RESULTS



Divya Rawat et. al. in preparation



- AstroSat and Insight-HXMT provide wide band spectra and simultaneous rapid timing information
- This allows for correlating QPO frequency with disk parameters ( inner radius and accretion rate)



## Summary....

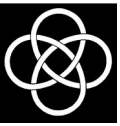
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- Using BOTH AstroSat and Insight-HXMT, for GRS 1915+105, the C-type QPO frequency is found to be a function of accretion rate and inner radii, thus identifying the frequency as the General Relativistic Dynamic (sound crossing time) one as predicted 49 years ago.
- Need to confirm using other data sets and for other sources..



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THANK YOU



# Spectral model dependency

